

Geometry of the LMC Disk: Results from MACHO and 2MASS

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Abstract

Wehavecross-correlatedMACHOLMCCepheidswith2MASSSecondIncrementalRelease Catalog. TheresultingdatabaseisconsiderablylargerthanthesetofOGLECepheidsinthe LMCbar, and has significantly better are alcoverage, allowing more accurate determination of LMC geometry. Random-phase correction is applied to 2MASSJ, H, and K smagnitudes, using the knowledge of V-bandlight curve and the ephemeris of 2MASS observations, to produce mean magnitudes. The improvement of phase-corrected PL relations over random-phase PL relations is clearly demonstrated. Reddening is estimated for each star individually, further

improving the accuracy of the method. The orientation parameters of the LMC are derived by a Maximum Likelihood approach which solves for viewing angles and PL coefficients simultaneously, providing an unbiased estimation. The results of the analysis are used to place limits on warping of the LMC disk. Implications for the microlensing optical depth are also discussed.

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